"COHERENT+ decay-at-rest coherent v"

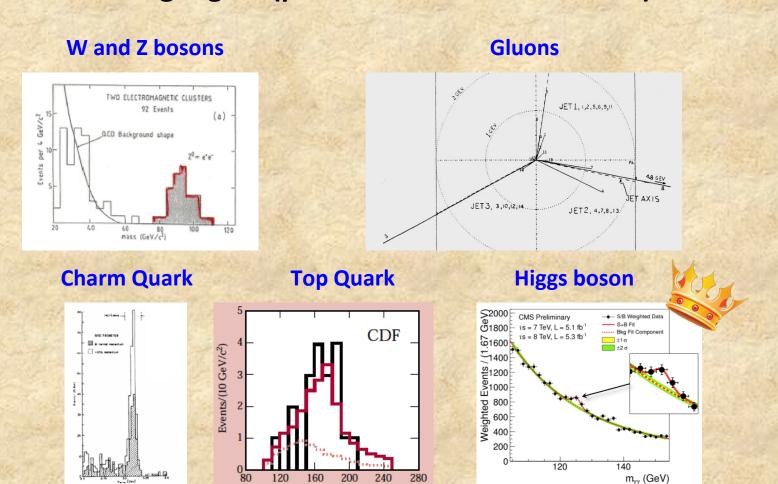
Yuri Efremenko University of Tennessee

WINP Feb 5, 2015

Standard Model

Formulated ~ 50 years ago, become driving force for HEP ever since

Short list of highlights (predictions and discoveries):

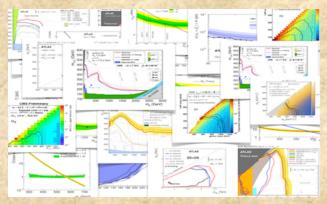


Search for extension of the SM

It is motivated by unexplained phenomena like Dark Mater and Dark Energy, Gravity, Neutrino masses, matter-antimatter asymmetry

Some big price tag examples

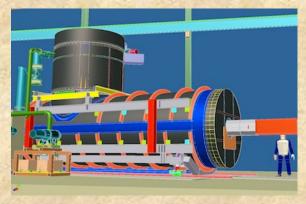
Search for SUSY at LHC



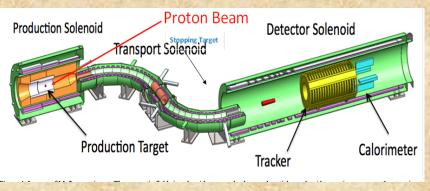




Search for Neutron EDM



mu2e



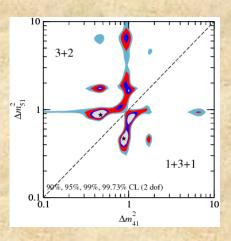
Do we have smoking guns in neutrino sector?

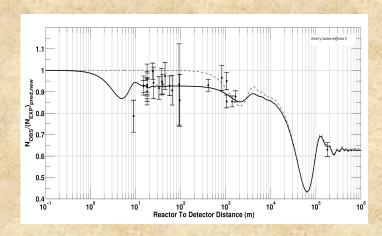
Non zero neutrino mass and large amplitude mixing are well established facts

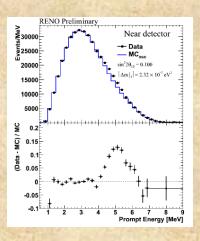
Sterile Neutrinos

Reactor Anomaly

Bump in reactor spectra







There is a world wide effort to confirm or rule them out

Are we extensively checking neutrino sector to look for the vSM?

Neutrino Neutral Current Coherent

Scattering

$$\frac{d\sigma}{dT_A} = \frac{G_F^2}{4\pi} m_A \left[Z \left(1 - 4\sin^2 \theta_W \right) - N \right]^2 \left[1 - m_A \frac{T_A}{2E_v^2} \right] F^2(Q^2)$$

$$\sigma_{tot} = \frac{G_F^2 E_v^2}{4\pi} \Big[Z \Big(1 - 4\sin^2 \theta_W \Big) - N \Big]^2 F^2(Q^2)$$

 m_A – nucleus mass

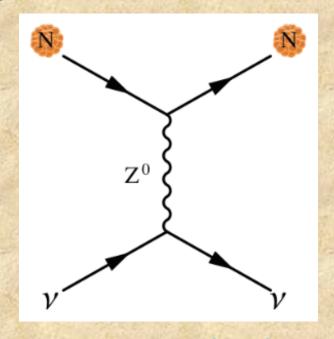
 T_A – kinetic energy of recoil nucleus

 E_{v} – neutrino energy

Z – nucleus charge

N – number of neutrons in the nucleus

F is nucleus form factor



D.Z. Freedman PRD 9 (1974)

A. Drukier & L. Stodolsky, PRD 30, 2295 (1984)

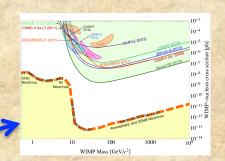
Horowitz et al. astro-ph/0302071

$$E_{\nu} < 50 MeV$$

This process has well predicted cross section.

It never been detected.

It is playing important role in core collapse SN dynamics
Could become irreducible background for DM experiments



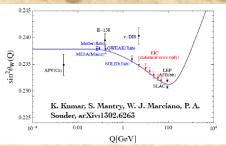
If we detect it and see agreement with the SM

Next after the "first light" → experiments with multiple targets

Tool for precise measurement of electro-week angle at low Q²

$$\sigma_{tot} \approx \left[Z \left(1 - 4 \sin^2 \theta_W \right) - N \right]^2$$

Require a few different targets

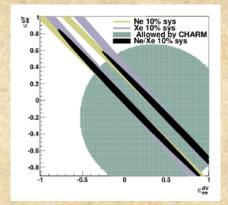


Search for nonstandard interactions K.Scholberg arXiv 0511-042

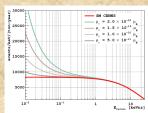
$$\frac{d\sigma}{dT_{coh}} = \frac{G_f^2 M}{2\pi} ((G_V + G_A)^2 + (G_V - G_A)^2 (1 - \frac{T}{E_\nu})^2 - (G_V^2 - G_A^2) \frac{MT}{E_\nu^2})$$

$$G_V = ((g_v^p + 2\epsilon_{ee}^{uV} + \epsilon_{ee}^{dV})Z + (g_v^n + \epsilon_{ee}^{uV} + 2\epsilon_{ee}^{dV})N)F_{nucl}^V(Q^2)$$

$$G_A = ((g_a^p + 2\epsilon_{ee}^{uA} + \epsilon_{ee}^{dA})(Z_+ - Z_-) + (g_a^n + \epsilon_{ee}^{uA} + 2\epsilon_{ee}^{dA})(N_+ - N_-))F_{nucl}^A(Q^2)$$



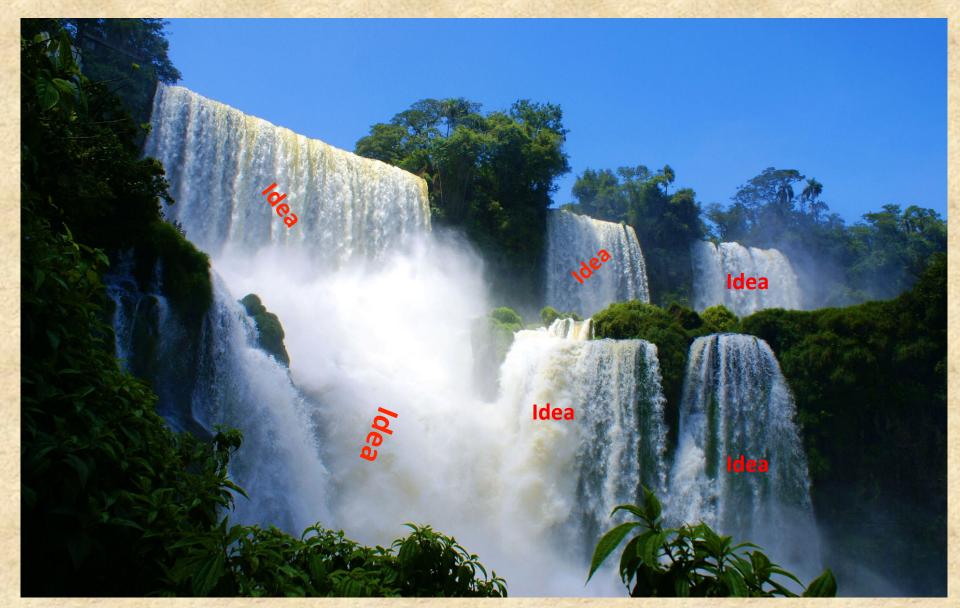
Search for Neutrino Magnetic Moment Brice, S.J. et al. Phys.Rev. D89 (2014)



There is a big interest from the theory community.

Please see "Coherent Theory Workshop, Raleigh, NC, January 2015, 30 attendees" (http://coherent-theory.phy.duke.edu/)

If we detect it and do not see agreement with the SM – waterfall of ideas will follow



What is the right place to look for Neutrino coherent scattering?

Nuclear reactor

or

DAR facility





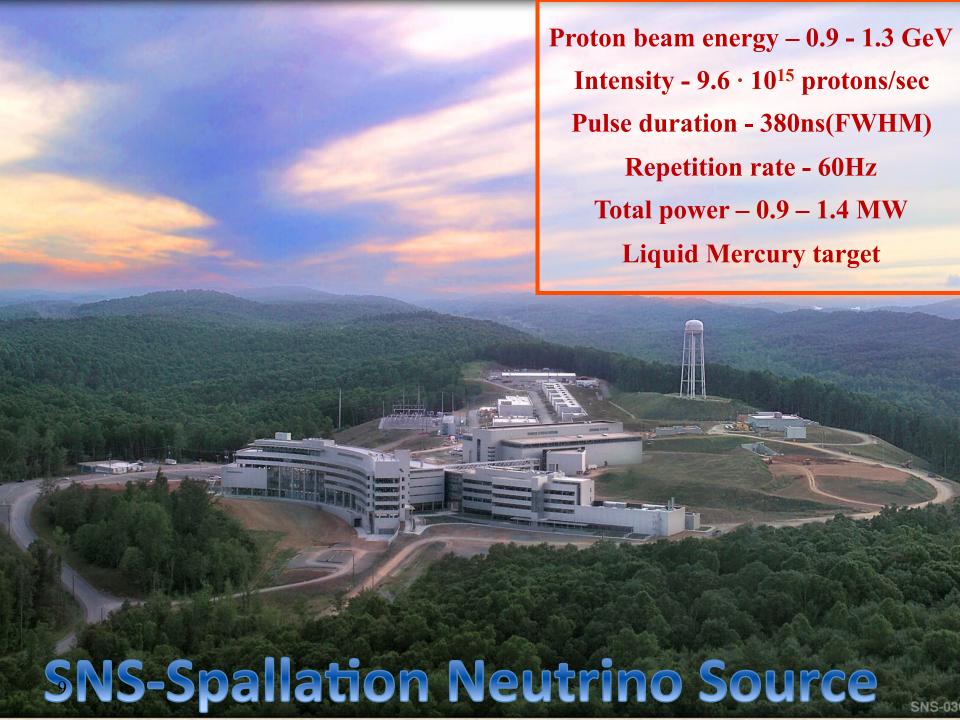
Huge neutrino flux

- Just right energy range
 - Pulsed structure
- Characteristic time distribution
 - Multiple flavors
 - Wide mass range of targets

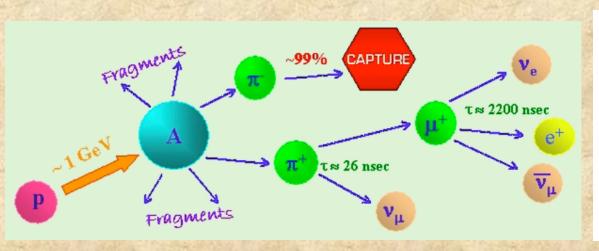


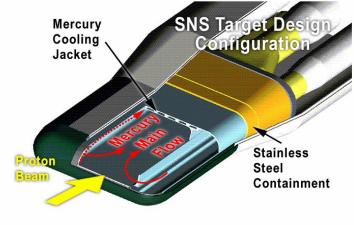
- Low energy: difficult to use heavy targets
 - No pulsed structure

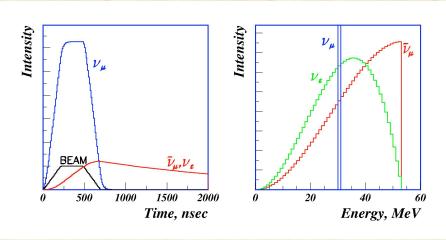
Not as large flux as at reactor



Neutrino Production at SNS







At present power SNS generates:

~5•10²⁰ protons per day,

or ~1•10²⁰ neutrinos of 3 flavors per day.

No upgrade or changes in the target are required



Collaboration to make the first detection of the Neutrino Neutral Current Coherent scattering at the SNS





University of California, Berkeley

University of Chicago

Duke University

University of Florida

Indiana University

Institute for Theoretical and Experimental Physics, Moscow

Lawrence Berkeley National Laboratory

Los Alamos National Laboratory

National Research Nuclear University MEPhl

North Carolina Central University

Oak Ridge National Laboratory

Pacific Northwest National Laboratory

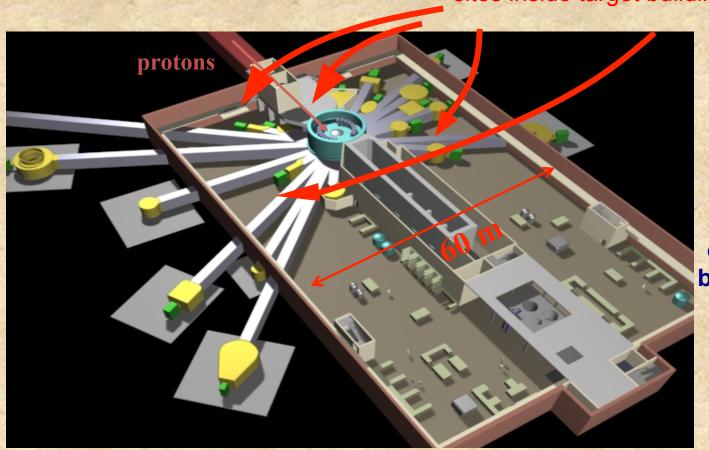
Sandia National Laboratory

University of Tennessee, Knoxville

Triangle Universities Nuclear Laboratory

Potential Locations for Neutrino Experiment at the SNS

sites inside target building including basement



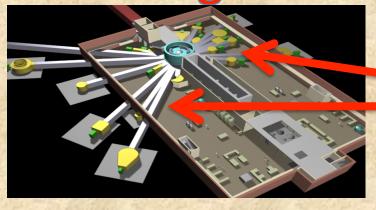
Multiple sites are available at a distance 15-20 m.

"Green field" is outside of the target building for distances more than 30 m

ORNL is strongly supporting BG studies for neutrino experiment at the SNS

ORNL support: 3 LDRD's (>\$300k) + Wigner Fellow

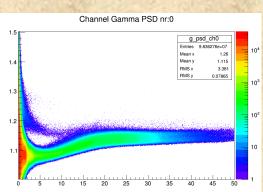
Background Measurements at SNS



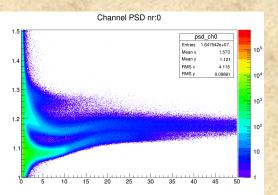


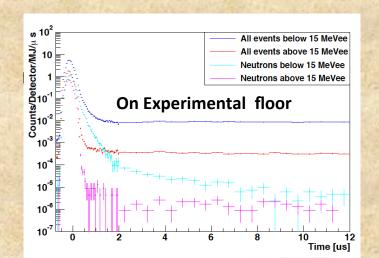
Started in Sept 2013
Tons of data at various locations

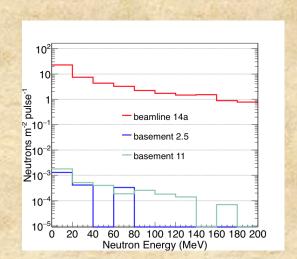
"Out-of-beam" events, primarily muons.



"In-Beam" events, considerably more neutron events





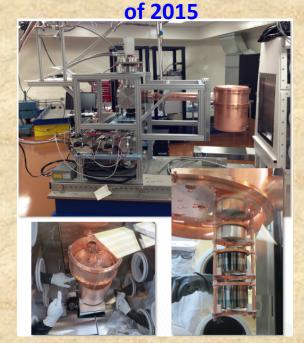


Three detector technologies are "Ready"

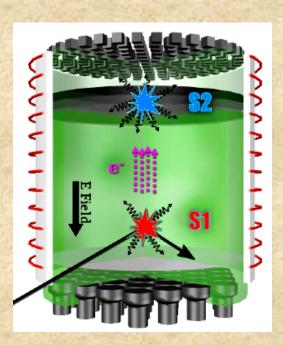
MJD prototype cryostat with 20 kg of HPGe detectors, could be available by the end

14 kg low background Csl crystal is available at the University of Chicago

100 kg, 2 phase LXe is detector being built at MEPhI, Moscow





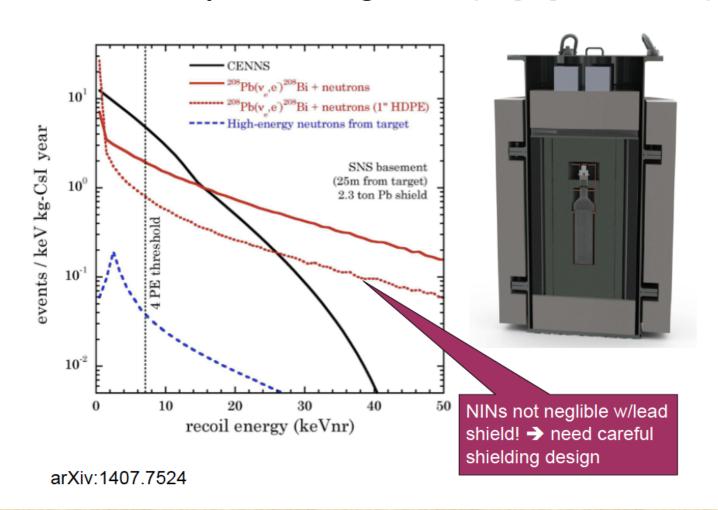


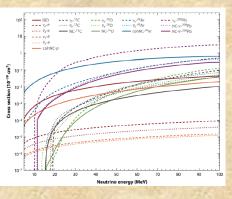
15 m from the target, 100 kg detector, prompt 30 MeV neutrinos

Target	Max Recoil (keV)	Cross section 10 ⁻⁴² cm ²	Threshold, keV _{nr}	N events, year
Ge	27	5830	3	2560
ı	15	19400	10	732
Xe	15	22300	1	5970

Realization that neutrinos can induce neutron background in the shielding

Estimate for a specific configuration (Csl[Na] in lead shield):





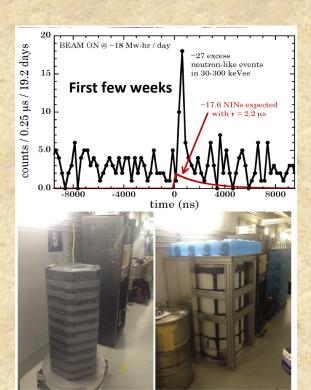
Neutrino Induced Neutrons (NIN)

Never been measured.

There are only theoretical calculations

This reaction on Lead is used by HALO experiment in the SNOlab, to watch for supernovae.





Measurement of NIN → the first neutrino experiment at the SNS

Liquid Scintillator detectors inside Lead, Poly, Cd, Water shield with muon veto

Started in the fall of 2014. By the summer of 2015 should have reasonable statistics to see effect

Planning to measure as well NIN on Iron and Copper for input for Nuclear Theory and shielding optimizations



Conclusion



Neutrino Neutral Current coherent scattering can be measured

This program is pioneered in the U.S.

There is broad interest from test of the SM, to astrophysics, and DM searches

SNS DAR is very attractive place to do experiment

COHERENT collaboration has been created

Good support from ORNL

Detectors for the 'first light" are available (+ international contribution)

Only modest investment required

Fast turn around and possibility to have first physics results within a year

Another possible location for this program is at FNAL (see Rex Tayloe talk)